

Fun with Cyclic Spectroscopy

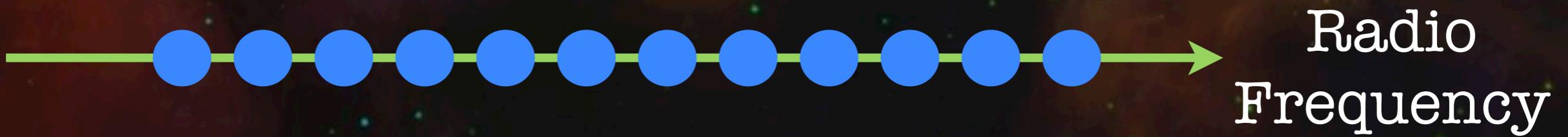
Mark Walker
(Manly Astrophysics)

Paul Demorest (NRAO)
Willem van Straten (Swinburne)
Aris Karastergiou (Oxford)

Overview

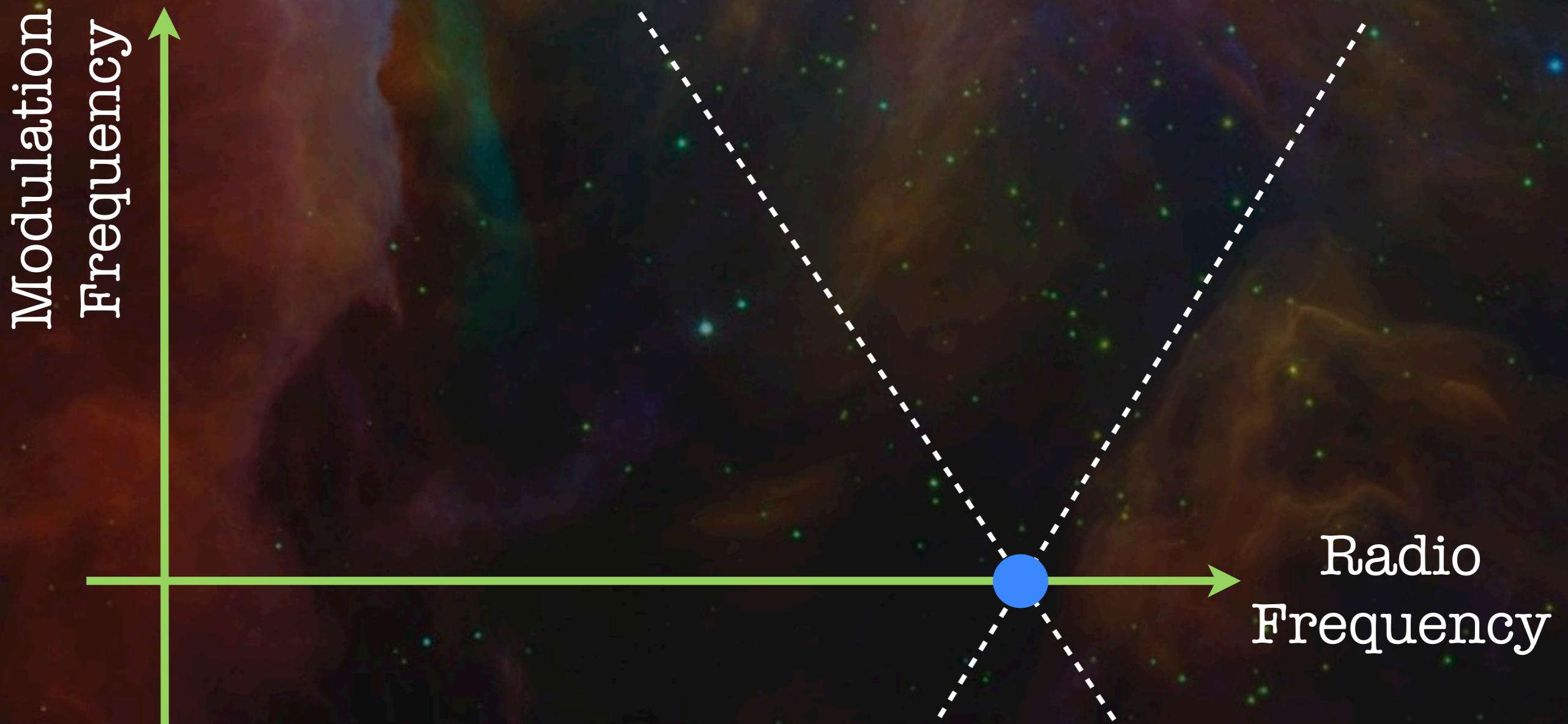
- What is cyclic spectroscopy ?
- Extracting information from cyclic spectra
- Holographic imaging
- Anticipated developments

Spectroscopy



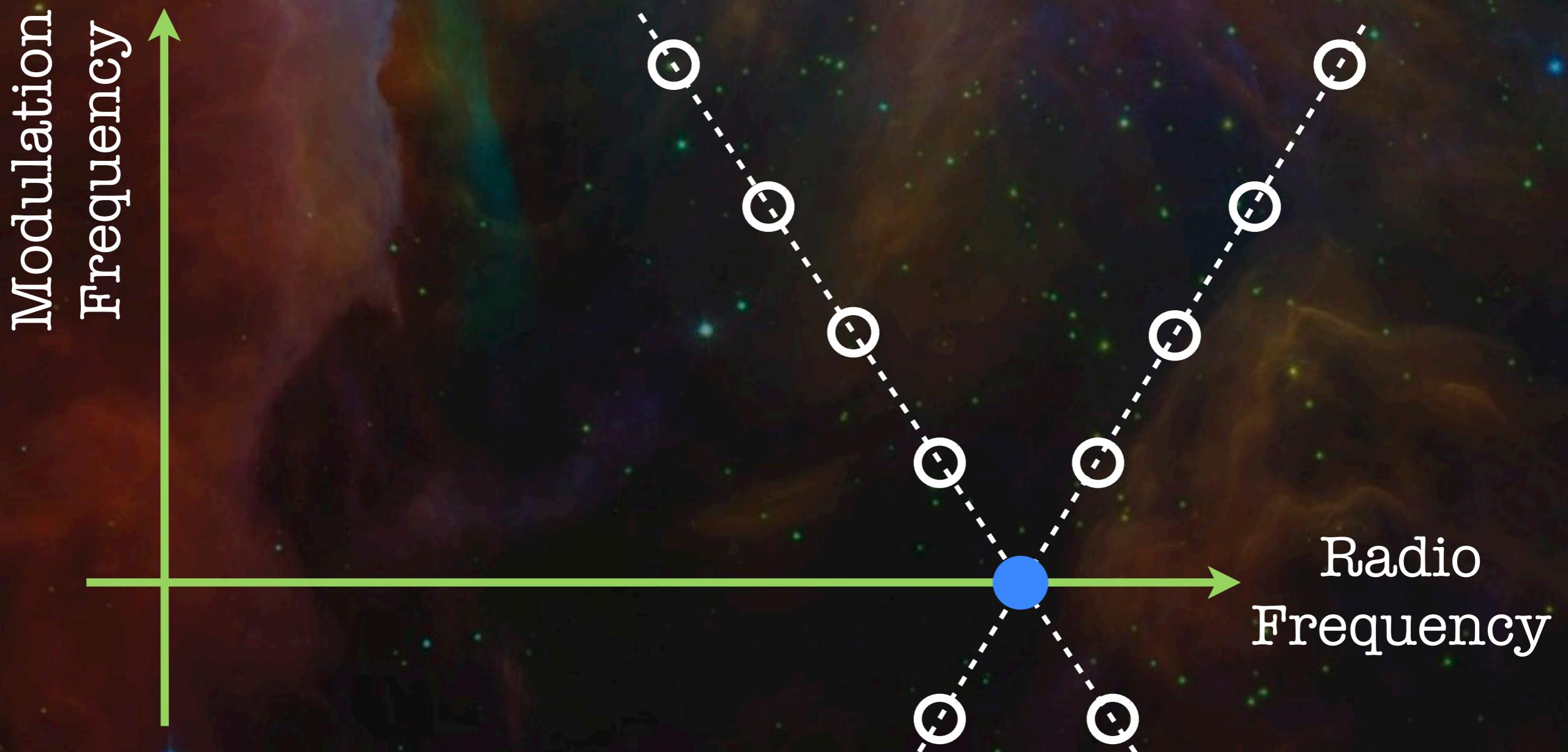
Cyclic Spectroscopy

Demorest 2011, MNRAS



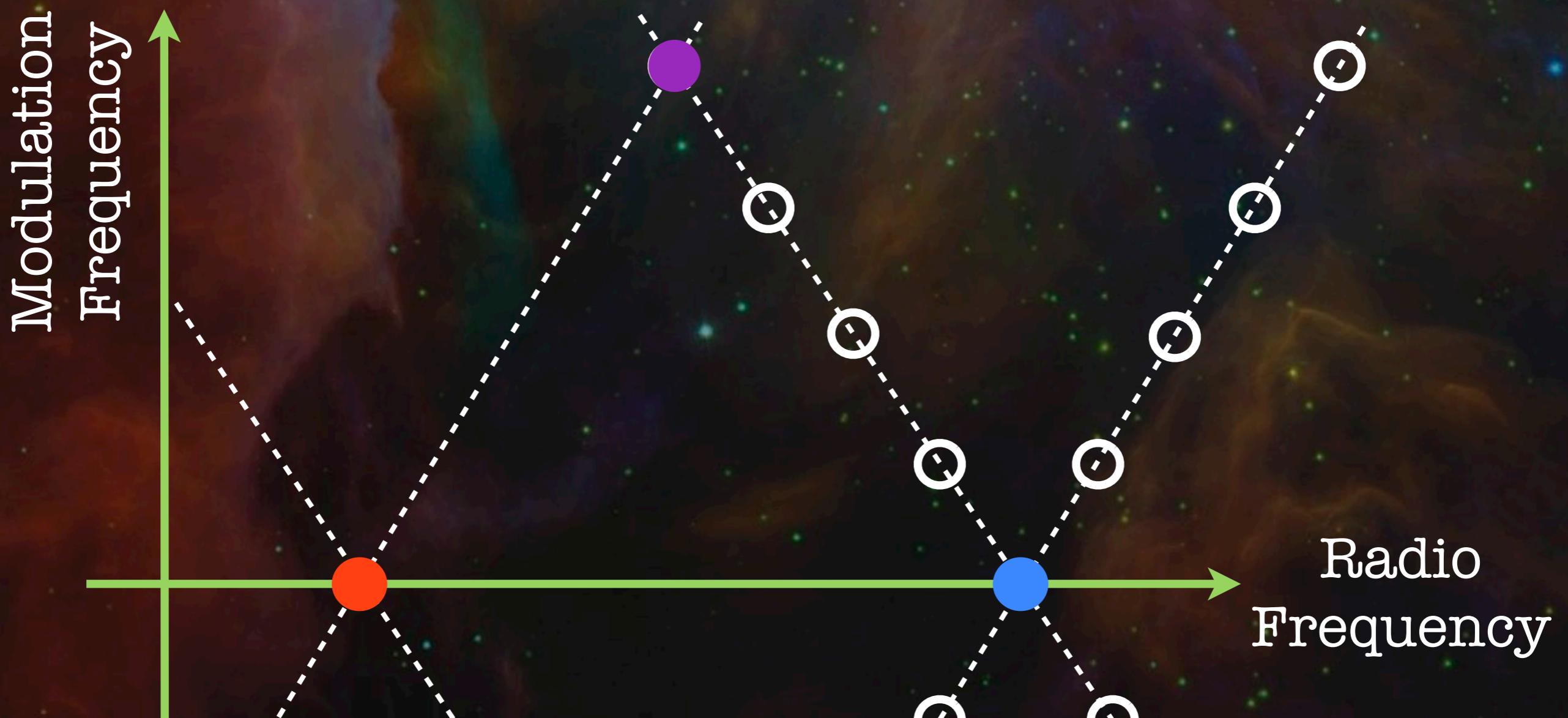
Cyclic Spectroscopy

Demorest 2011, MNRAS



Cyclic Spectroscopy

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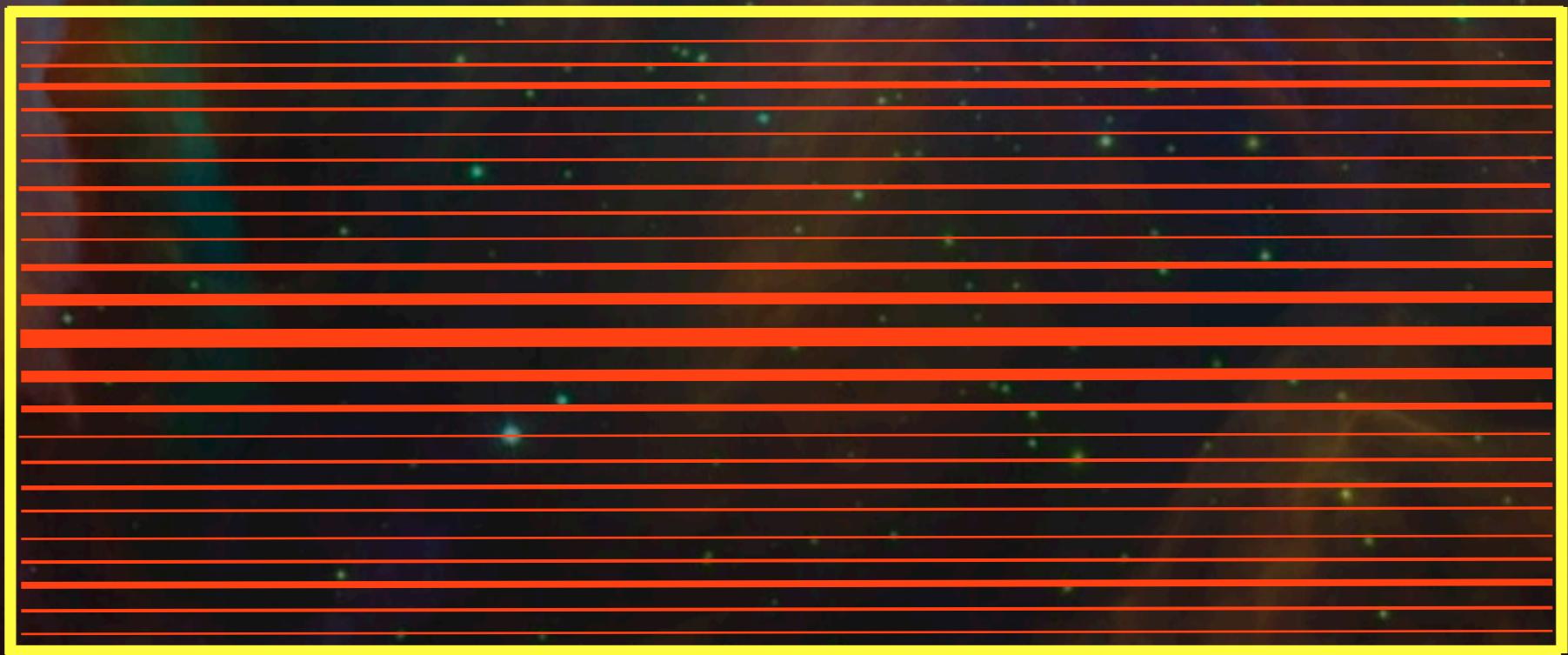


$$S(\alpha, \nu) = \langle E(\nu + \alpha/2) E^*(\nu - \alpha/2) \rangle$$

Intrinsic Cyclic Spectrum

Radio Frequency

Modulation
Frequency

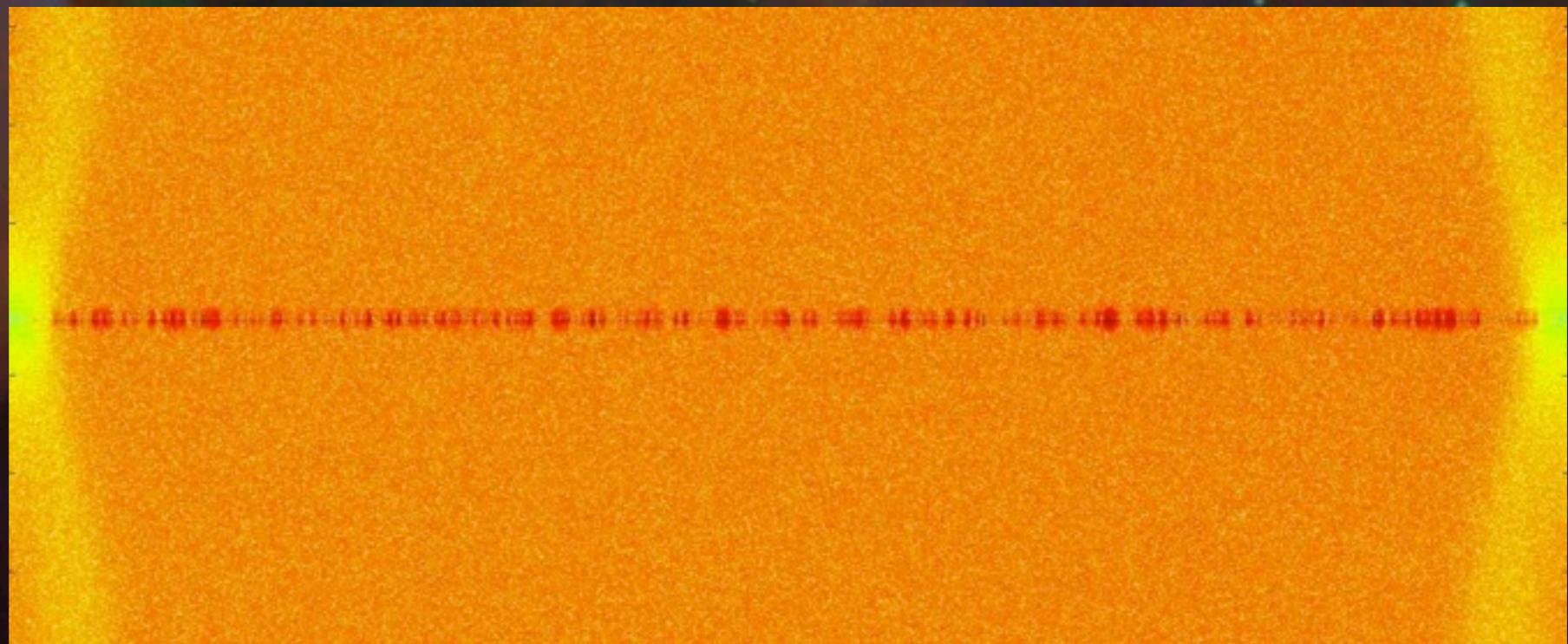


$$S_o(\alpha, \nu) = \langle E_o(\nu + \alpha/2) E_o^*(\nu - \alpha/2) \rangle \rightarrow S_o(\alpha)$$

B1937+21 @ Arecibo, 428 MHz

Radio Frequency

Modulation
Frequency

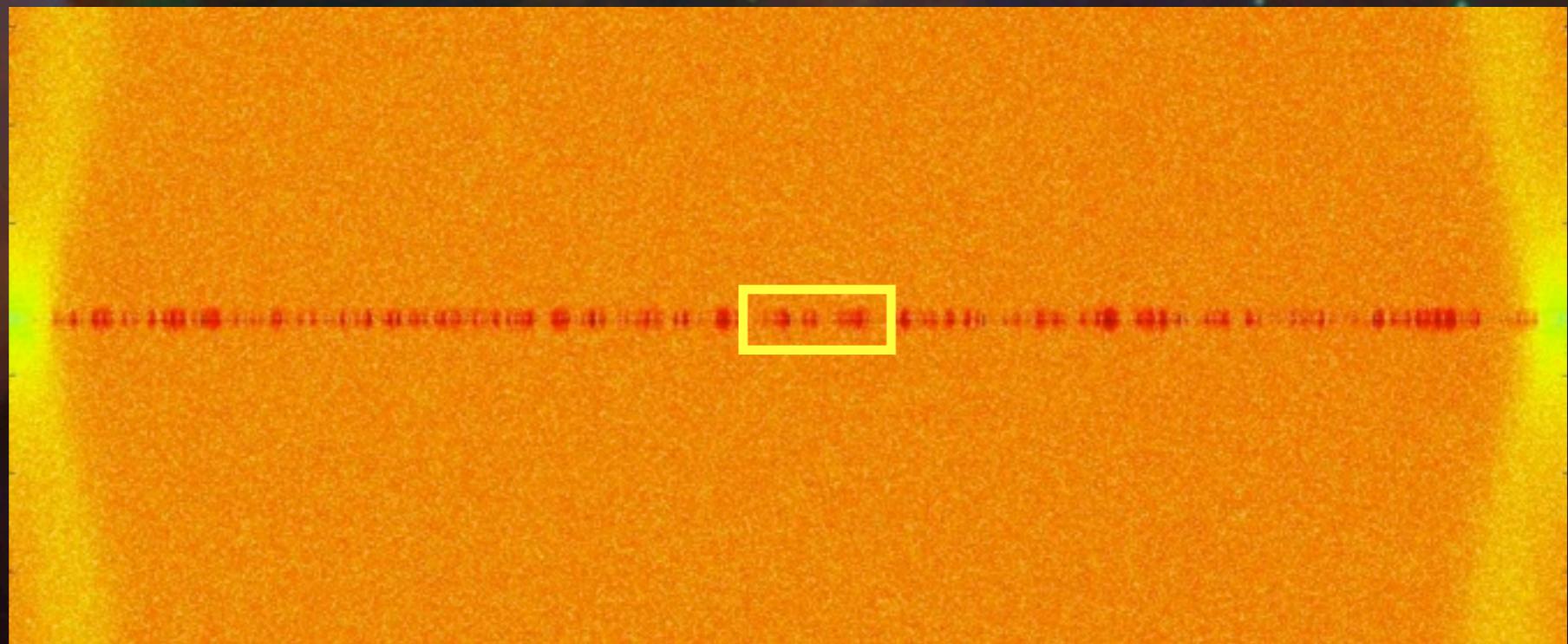


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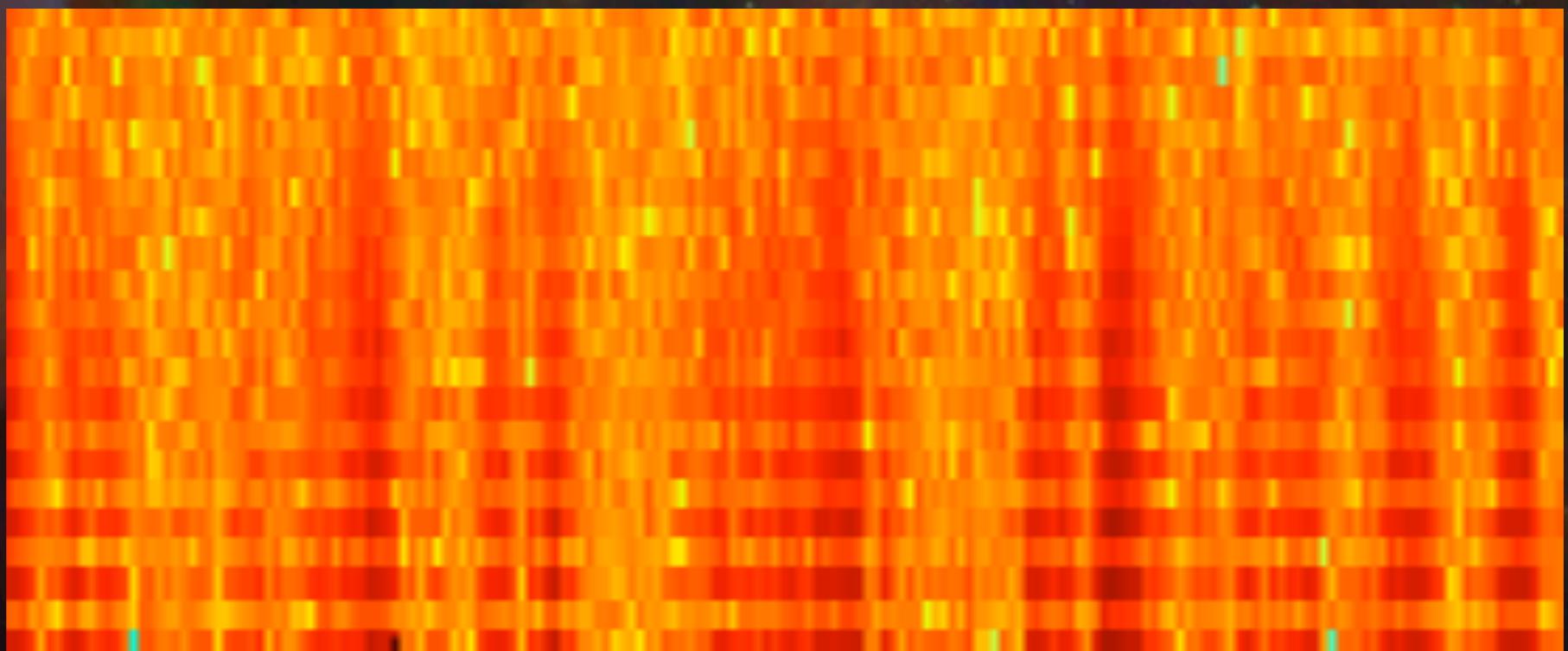


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$$S_o(\alpha, \nu) = \langle E_o(\nu + \alpha/2) E_o^*(\nu - \alpha/2) \rangle \rightarrow S_o(\alpha)$$

$$E(\nu) = H(\nu) E_o(\nu)$$

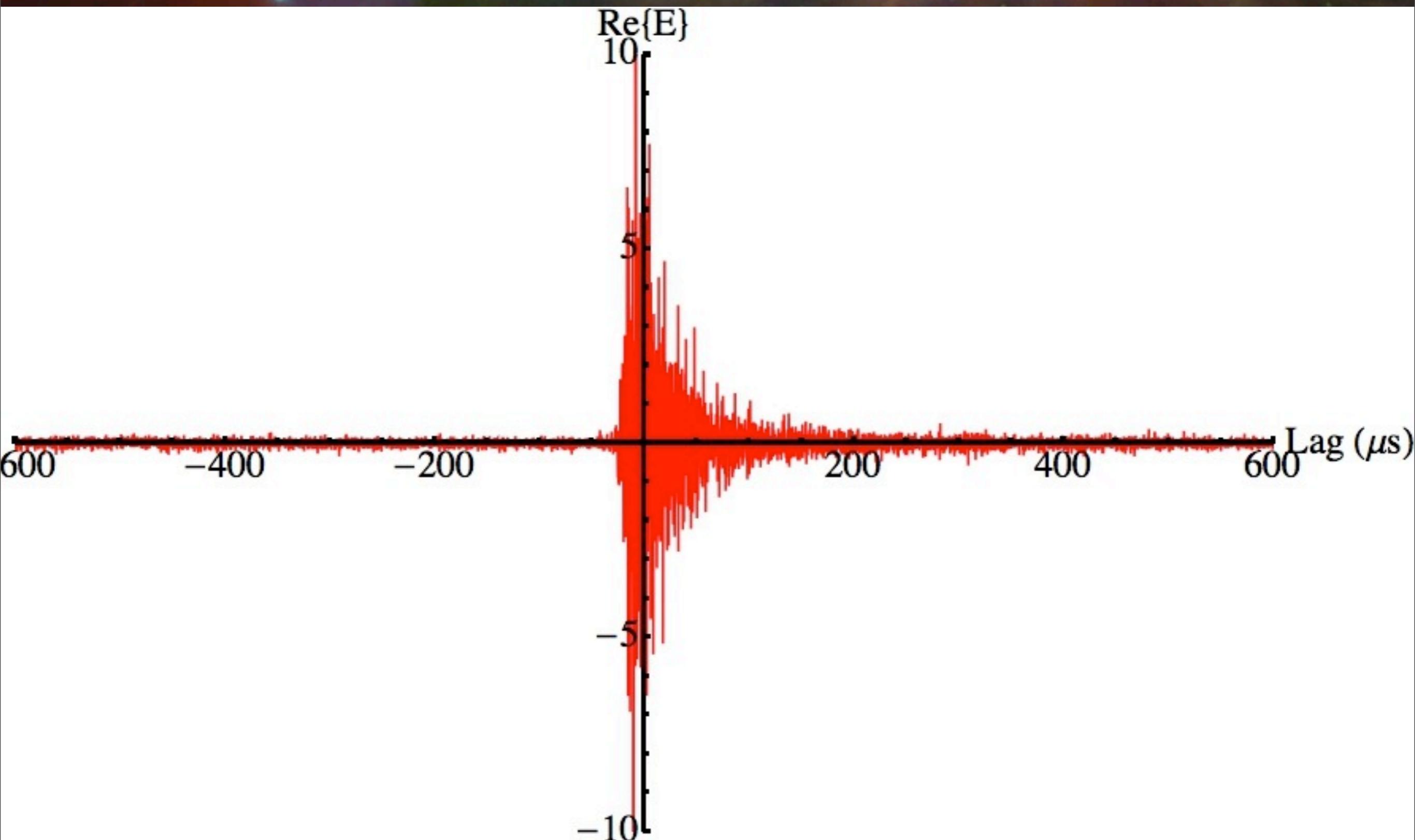
$$S(\alpha, \nu) = H(\nu + \alpha/2) H^*(\nu - \alpha/2) S_o(\alpha)$$

Extracting $H(\nu)$ and $S_o(\alpha)$

$$S(\alpha, \nu) = H(\nu + \alpha/2) H^*(\nu - \alpha/2) S_o(\alpha)$$

- Given $S_o(\alpha)$ we can fit for $H(\nu)$
- Least-squares fit for $h(\tau)$ using L-BFGS

Extracting $H(v)$ and $S_o(\alpha)$

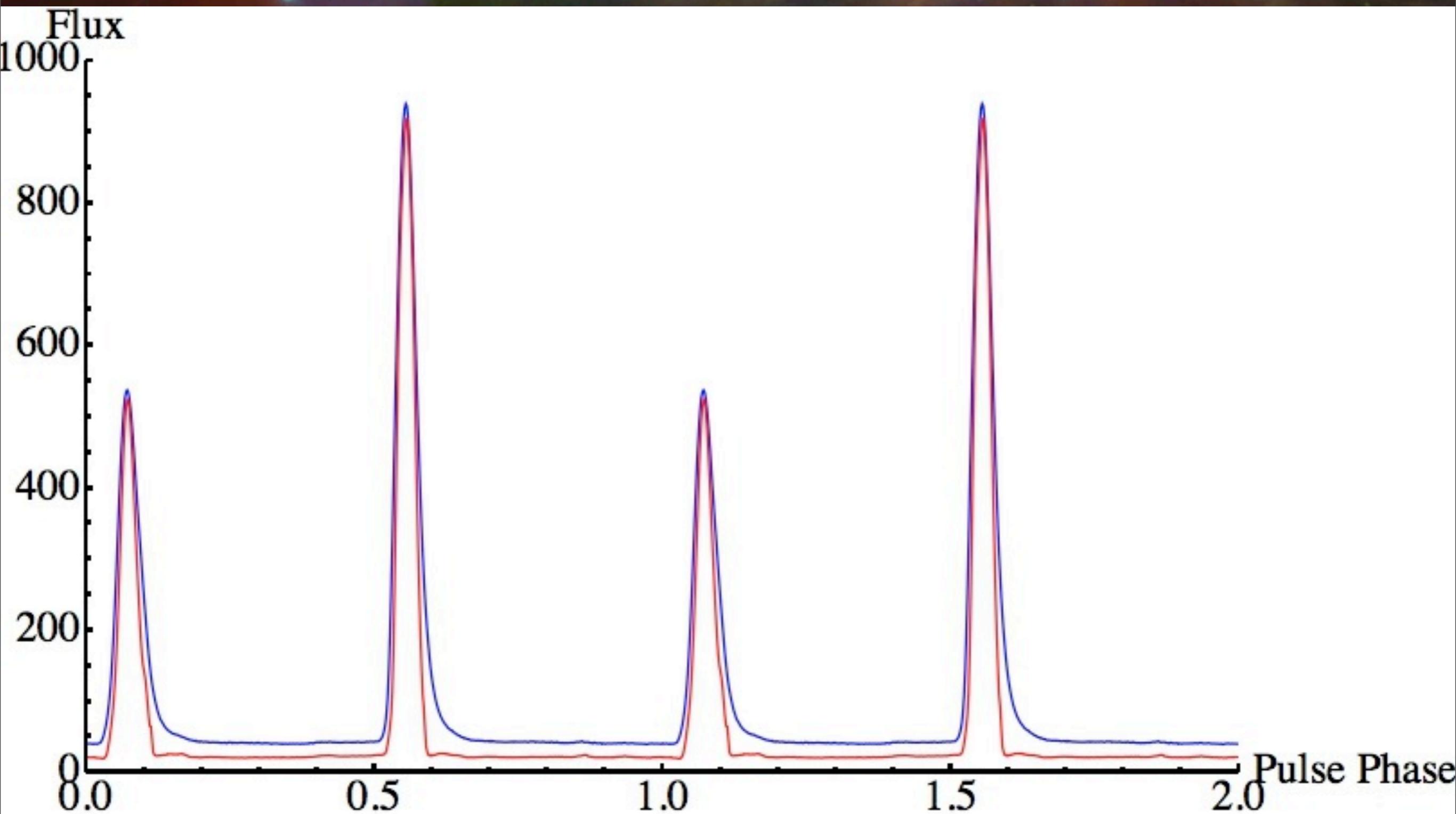


Extracting $H(\nu)$ and $S_o(\alpha)$

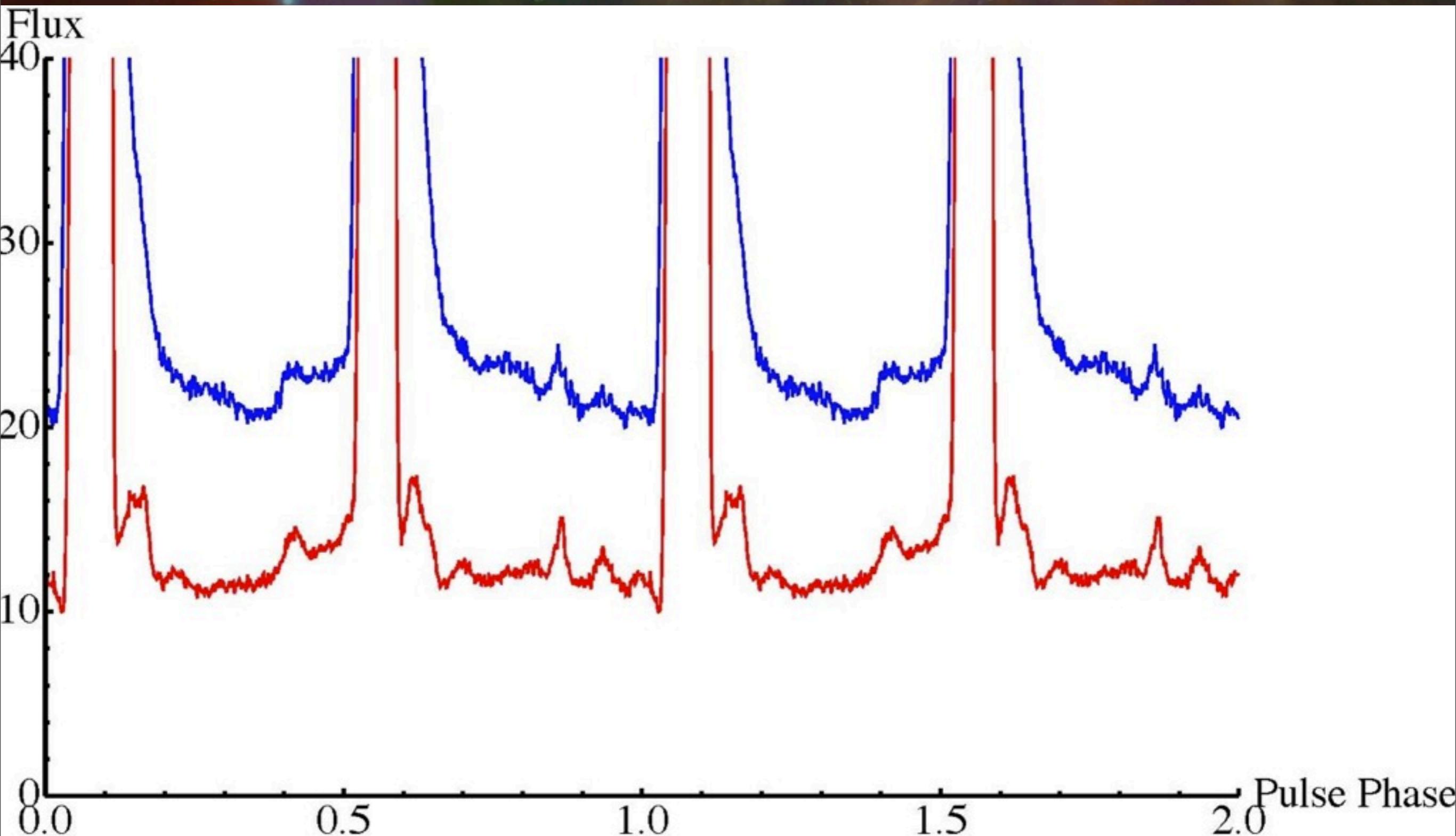
$$S(\alpha, \nu) = H(\nu + \alpha/2) H^*(\nu - \alpha/2) S_o(\alpha)$$

- Given $S_o(\alpha)$ we can fit for $H(\nu)$
 - Least-squares fit for $h(\tau)$ using L-BFGS
- Given $H(\nu)$ we can fit for $S_o(\alpha)$
 - One-step least-squares fit
- Average $S_o(\alpha)$ over whole observation
- Iterate

Extracting $H(\nu)$ and $S_o(\alpha)$



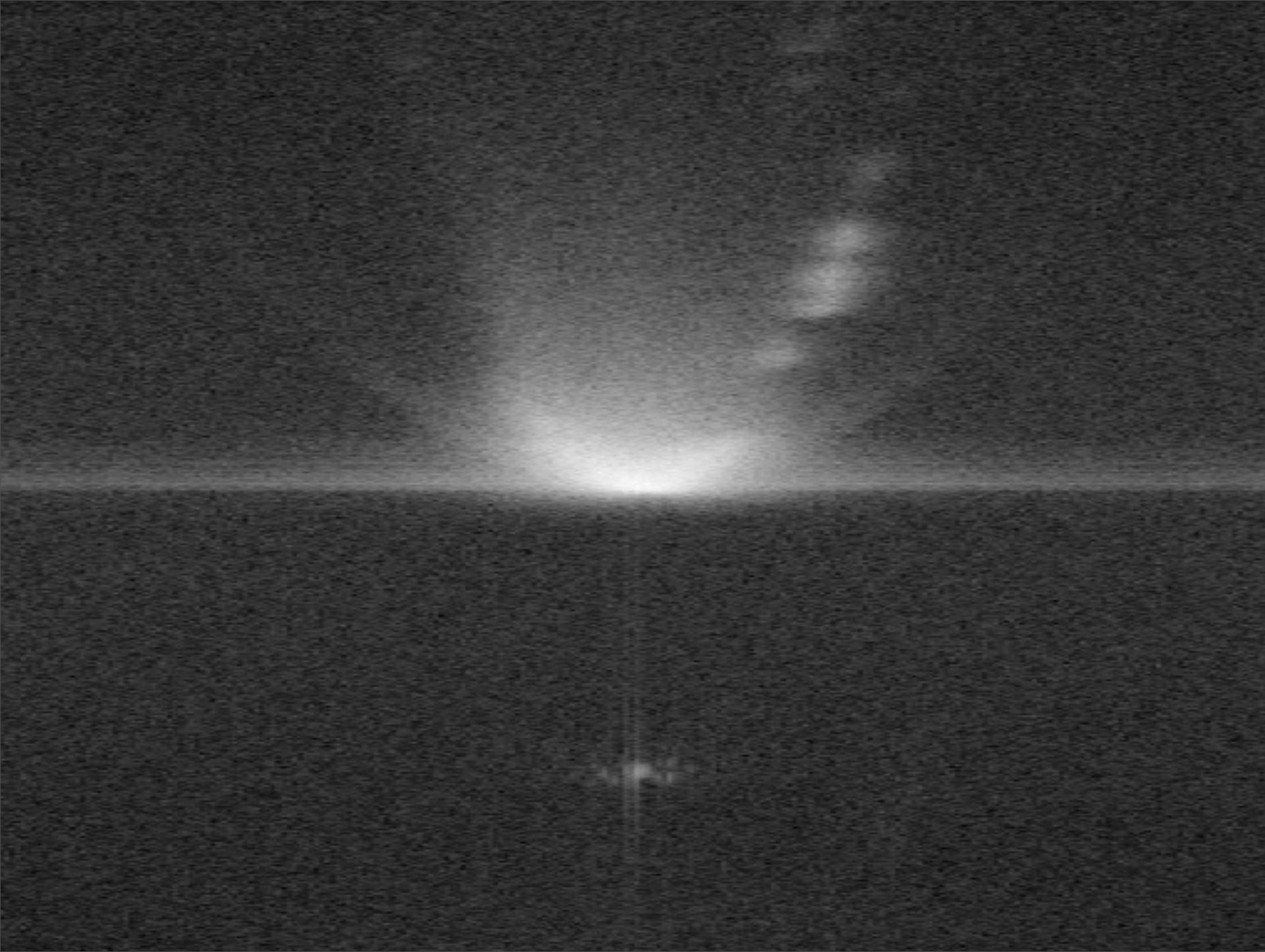
Extracting $H(\nu)$ and $S_o(\alpha)$



Extracting $H(\nu)$ and $S_o(\alpha)$

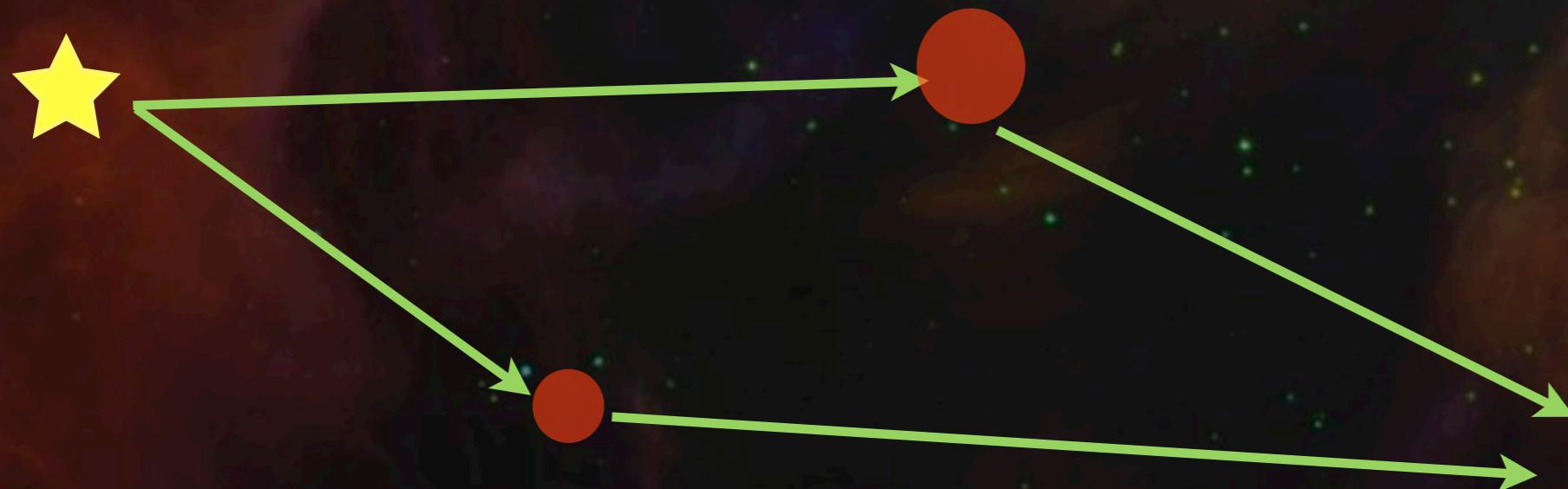
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- Average $S_o(\alpha)$ over whole observation
- Iterate
- $H(\nu)$ for every 10 sec sample $\rightarrow H(\nu, t)$
- Hence $h(\tau, \omega)$

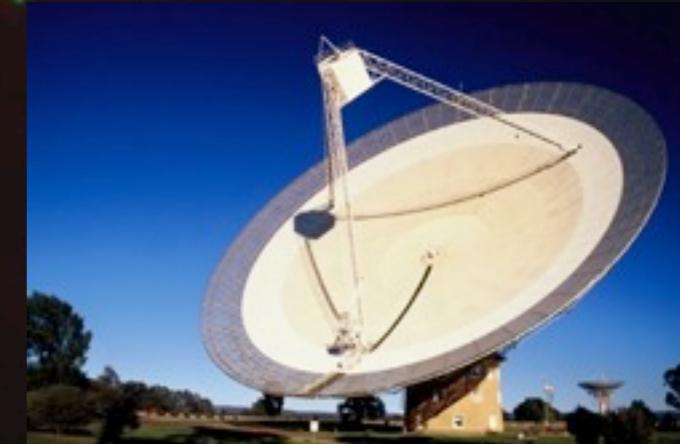


Holographic Imaging

- Have detailed holographic information directly from cyclic spectrum : $h(\tau, \omega)$
- Need to model this as $n_e(x, y, z)$ on line-of-sight
 - Physics of ISM on AU-scales; Accurate TOAs
 - Accurate pulse timing



- GBT observations of J1518+4904



Anticipated Developments

- Currently only Stokes-I : want full Stokes
- Currently produce CS by post-processing voltages
 - Want real-time products :
 - $V(t) \rightarrow \text{Stokes}(\alpha, \nu) \rightarrow h(\tau) \text{ and } S_o(\alpha)$
- New Arecibo & GBT FPGA-based backends
 - Glenn Jones (Caltech) & PD
 - ~ 200 MHz (feeding PUPPI/GUPPI cluster)
- PKS : HIPSR ~ 400 MHz ?